



PROGRESS REPORT COVERING THE PERIOD MARCH 1, 1966 TO AUGUST 31, 1966.

FOR GRANT NO. 21-002-059

to

NATIONAL AERONAUTICS AND SPACE ADMINISTRATION

from

DEPARTMENT OF CHEMISTRY UNIVERSITY OF MARYLAND COLLEGE PARK, MARYLAND

in cooperation with

NATIONAL BIOMEDICAL RESEARCH FOUNDATION SILVER SPRING, MARYLAND

CFSTI PRICE(S) \$

Hard copy (HC) // / C

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October 1966

The following progress has been made on Grant No. 21-002-059, covering the period March 1, 1966 to August 31, 1966.

Gas mixtures representing a wide range of C:H ratios have been subjected to the plasma discharge and the product mixtures are being It has been found by gas chromatography that they contain a large number of polynuclear aromatic hydrocarbons in small quantities. To aid in the identification of these compounds, most of the commerically available unsubstituted polynuclear aromatic hydrocarbons have been run in the gas chromatograph under the conditions best suited for separation of the product mixtures. It was found that the log of the relative retention times gives a straight line when plotted against the boiling points, and a smooth curve when plotted against the number of carbon atoms. These plots, together with chromatograms of the product mixtures, afford some information on reaction products for which no standards are available for direct comparison. A number of specific compounds have been identified.

A number of theoretical calculations have been carried out to serve as an aid in evaluating the experimental results.

The following table gives computed Equilibrium Concentrations of Selected Compounds which, under certain conditions, should be present in greater than 1% mole fractions.

Temperature	Pressure	Aromatics Excluded	Aromatics Included
300°K	10 ⁻⁶ Atm.	Methane H ₂ 1,3-Butadiene Cyclohexane	Asphalt Methane H ₂
	1 Atm.	Methane H ₂ Acetylene Cyclohexane 1,3-Butadiene Cyclohexene Methylcyclohexane	Asphalt Methane H ₂

Temperature	Pressure	Aromatics Excluded	Aromatics Included
300°K	100 Atm.	Methane H ₂ Acetylene Cyclohexane 1,3-Butadiene Cyclohexene Methylcyclohexane	Asphalt Methane H ₂
500°K	10 ⁻⁶ Atm.	Methane H ₂ Acetylene Ethylene Propyne 1,3-Butadiene Isoprene Cis-piperylene Trans-piperylene Cyclohexene	Asphalt Methane ^H 2
	1 Atm.	Methane H2 Acetylene Ethane Cyclopentane Cyclohexane Propyne 1,3-Butadiene Isoprene Cis-piperylene Trans-piperylene Cyclohexene Methylcyclohexene	Asphalt Methane ^H 2
	100 Atm.	Methane H ₂ Acetylene Cyclopentane Cyclohexane Propyne 1,3-Butadiene Isoprene Cis-piperylene Trans-piperylene Cyclohexene Methylcyclohexane	Asphalt Methane ^H 2
700°K	10 ⁻⁶ Atm.	Methane H ₂ Acetylene Ethylene Propyne	Asphalt Methane ^H 2

Temperature	Pressure	Aromatics Excluded	Aromatics Included
	1 Atm.	Methane H ₂ Acetylene Ethylene Ethane Propene Propyne Allene 1,3-Butadiene Isoprene Cis-piperylene Trans-piperylene Cyclohexene	Asphalt Methane ^H 2
	100 Atm.	Methane H ₂ Acetylene Ethane Propane Cyclopentane Propene Propyne Allene 1,3-Butadiene Isoprene Cis-piperylene Trans-piperylene Cyclohexene Methylcyclohexane	Naphthalene Asphalt Methane
1000°K	10 ⁻⁶ Atm.	H ₂ Acetylene	Asphalt H ₂ Acetylene
	1 Atm.	Methane H ₂ Acetylene Ethylene Propene Propyne Allene 1,3-Butadiene Isoprene Cis-piperylene Trans-piperylene	Benzene Naphthalene Asphalt Methane H ₂
	100 Atm.	Methane H ₂ Acetylene Ethylene Ethane Propene Butane	Benzene Naphthalene Asphalt Methane H ₂

Temperature Pressure Aromatics Excluded Aromatics Included

Propyne
Allene
1,3-Butadiene
1,4-Pentadiene
Isoprene
Cis-piperylene
Trans-piperylene
Cyclohexene

The preparation of a book entitled "Thermodynamic Equilibria in Pre-biological Atmospheres" is in its final stages and should be finished sometime during the Fall of 1966.

The following represents completed work, reported in part previously but now completed and published:

"Thermodynamic Equilibrium and the Inorganic Origin of Organic Compounds", R.V. Eck, E.R. Lippincott, M.O. Dayhoff, Y.T. Pratt, Science, 153, 3736, 628 (1966).

Theoretical and experimental support is presented for the hypothesis that many organic compounds may form under conditions of thermodynamic equilibrium. This possibility must be considered along with special effects of selective catalysts, radiation, and degradation from biological matter, in explaining the origin of organic compounds in carbonaceous chondrites. Similar considerations may apply to solar The equilibrium distribution of nebulas and planetary atmospheres. organic compounds at temperatures between 300°K and 1000°K and pressures of 10⁻⁶ to 50 atm for the C—H—0 system have been computed. moderate temperatures and low pressures, conditions where graphite production is inhibited, aromatic compounds may form even in the presence of large excesses of hydrogen. Such conditions exist in the solar nebula and in the atmospheres of some of the major planets. concentrations of a large number of compounds at 1000°K with nitrogen. sulfur, and chlorine added to the system have also been determined.

In some cases, a limited equilibrium method is employed in which these few compounds which form with the most difficulty are excluded from the computations, while representatives of all other families of compounds are included. This approach is shown to the useful in the interpretation of certain experimental data in which complete equilibrium has not been attained. We have also found that gases, activated to the plasma state by a high-energy radio frequency field, recombine on cooling to yield product mixtures which are in qualitative agreement with those predicted by the equilibrium computations. We believe that such products can be profitably studied as if at a metastable limited equilibrium.